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High Spatial Resolution Thermal Satellite Technologies

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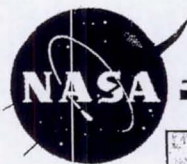
April 22-23, 2002³
Arizona State University



High Spatial Resolution Thermal Band

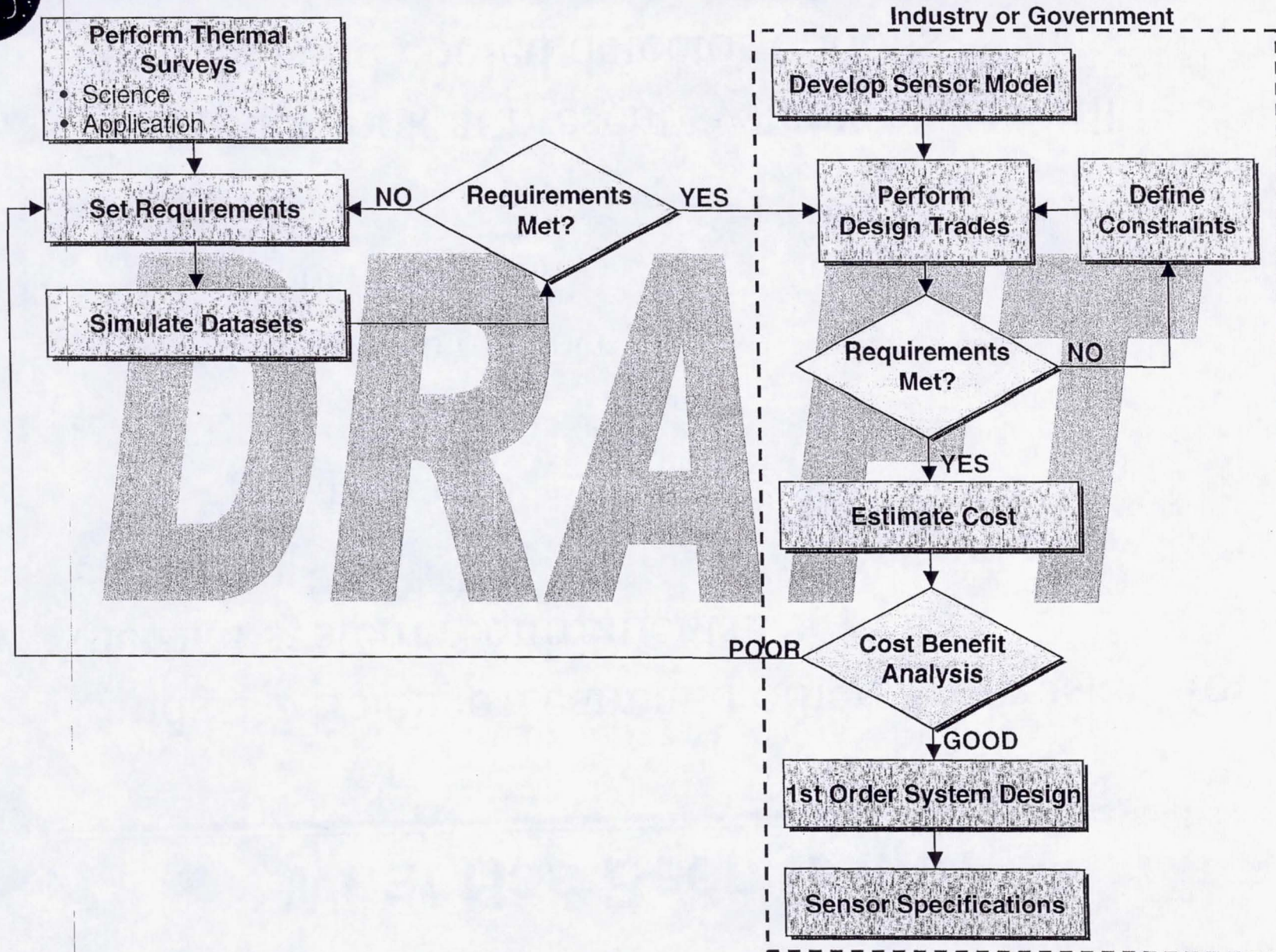
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- Problem
 - No follow-on exists for Landsat 7 or ASTER high spatial resolution thermal systems
 - Can new technologies be implemented to allow for a low-cost, useful alternative?
- Approach
 - Investigate both traditional cooled cross-track scanners and new architectures (cooled and uncooled) that could enable a low-cost thermal capability



General Approach

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What Has Been Done?

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- Landsat 7 Science Team and others were asked for thermal system requirements:
 - Spectral
 - Temperature range
 - NEDT
 - Ground sample distance
 - Point spread function
 - Revisit time
- First order look at present systems and possible uncooled and cooled detector options
 - Standard pushbroom options
 - Backscanned systems



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Research or Area of Interest	NEDT (K)	Tmin (K)	Tmax (K)	Radiometric Accuracy (%)	GSD (m)	PSF FWHM (m)	Revisit Time (days)	Spectral Band(s)	Other Requirements	Source
Cloud Detection for LTAP	.5K	198K	310K	5%	60 (ideal) 120 (acceptable)	76 m	per image			GSFC
Monitoring Surface Energy and Water Fluxes 1	~0.2	273	310	1k	~100	~100	~15	1-ok, 2 or more better	day/night pairs are very useful	RIT
Monitoring Surface Energy and Water Fluxes	0.1	263	343	0.5	20	40	7 (geostationary is better)	8-14 um, 6 bands equally spaced		JPL
Monitoring Surface Energy and Water Fluxes	0.2 K	250	350	3K (4%)	30m	30m	weekly	10-11microns	10-11AM local time overpass	UW-Madison
Material Transport in Aquatic Systems	~0.3	273	305	1 K Absolute	~100	~100	~15	1-ok, 2 to improve ATM cal would be better		RIT
Material Transport in Aquatic Systems 1	0.25	260	320	0.2 Absolute	60	ETM+	Hourly	Vis/TIR Split Window for ATM		Brown University
Agriculture Studies										JPL
Soil Moisture	0.1	273	313	0.5	20	40	7	same as above		JPL
Soil Moisture 1	0.2 K	250	350	3K (4%)	30m	30m	weekly	10-11microns		UW-Madison
Canopy Heat Exchange	0.1	263	343	0.5	20	40	7	same as above		JPL
Canopy Heat Exchange 1	0.2 K	250	350	3K (4%)	30m	30m	weekly	10-11microns		UW-Madison
Crop Health	0.1	263	313	0.5	20	40	7	same as above		JPL
Crop Health 1	0.2 K	250	350	3K (4%)	30m	30m	weekly	10-11microns		UW-Madison
Irrigated Crops	0.1	273	313	0.5	20	40	7	same as above		JPL
Irrigated Crops 1	0.5	273	313	Relative is important	120	120		ETM+ _		CU-Boulder
Irrigated Crops 2	0.2 K	250	350	3K (4%)	30m	30m	weekly	10-11microns		UW-Madison
	GSD Ground Sample Distance									
	PSF FWHM Point Spread Function Full Width Half Maximum (system blur)									
	Tmin, Tmax Temperature Range of Phenomena									
	NEDT Temperature Sensitivity									

[illegible]



Thermal Survey Summary

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- Many Landsat thermal applications will work with >60 meter GSD data
- Wide range of requirements indicates that more work is necessary to define them better
- This workshop should help provide such definition



Traditional TIR Architectures

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- Cross-track scanning systems
- Cooled either actively or passively
- Small number of HgCdTe detectors
- Typically large GSD; Landsat is the smallest of the cross-track scanning systems at 60 m
- Telescope diameter typically driven by SNR considerations and not diffraction
- Most systems are multispectral



Existing TIR Systems

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Satellite	Sensor	Architecture	Band	Spectral Range	Detector Material	Telescope Diameter [cm]	Spatial Resolution		Swath [km]	Quant. [bit]	NE Δ T [K]
				[μ m]			GSD [m]	MTF @ Nyquist			
Landsat 7	ETM+	Cross-track scanner	6	10.4 - 12.5	HgCdTe	40	60	0.3	185	8 of 9	
Terra	ASTER	Cross-track scanner	10	8.125 - 8.475	HgCdTe	24	90	0.35	60	12	0.2
			11	8.475 - 8.825			"		"	"	"
			12	8.925 - 9.275			"		"	"	"
			13	10.25 - 10.95			"		"	"	"
			14	10.95 - 11.65			"		"	"	"
NOAA	AVHRR	Cross-track scanner	4	10.3 - 11.3	HgCdTe	20	1100	0.3	2800	10	0.12
			5	11.5 - 12.5			"		"	"	"
Terra	MODIS	Cross-track scanner	27	6.535 - 6.895	HgCdTe	18	1000	0.35	2330	12	0.25
			28	7.175 - 7.475			"		"	"	0.25
			29	8.400 - 8.700			"		"	"	0.05
			30	9.580 - 9.880			"		"	"	0.25
			31	10.780 - 11.280			"		"	"	0.05
			32	11.770 - 12.270			"		"	"	0.05
			33	13.185 - 13.485			"		"	"	0.25
			34	13.485 - 13.785			"		"	"	0.25
			35	13.785 - 14.085			"		"	"	0.25
			36	14.085 - 14.385			"		"	"	0.35
MTI	MTI	Pushbroom	L	8.00 - 8.40	HgCdTe	36	20		12	12	0.025
			M	8.40 - 8.85			"		"	"	0.027
			N	10.20 - 10.70			"		"	"	0.045
ERS-2	ATSR-2	Conical scanner	1	11.5 - 12.3		11	1100		500	12	0.02
			2	10.6 - 11.3			"		"	11	"



Uncooled Pushbroom Pathfinding Thermal Instruments

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- Thermal Emission Imaging System IR Sensor (THEMIS)
 - Flown on Mars Orbiter launched April 2001 to map Martian surface mineralogy
 - Uncooled Raytheon microbolometer array (320x240 50 μm pitch pixels)
 - Pushbroom imager with precision-aligned stripe filter
 - 9 bands between 6.2 and 15.5 μm
 - f/1.6 optics
 - 4.4 degree FOV
 - 12.9 cm aperture
 - 30 Hz readout
 - 100 m GSD
 - ~ \$12M for instrument



New Thermal Architectures

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- Pushbroom systems and framing cameras can provide significant sensitivity advantages over cross-track scanners
- $\text{SNR} \sim (\text{no. of detectors})^{0.5}$
- Benefits of increased integration time associated with pushbroom systems and framing cameras:
 - Higher SNR
 - Smaller GSD
 - Potential use of uncooled detectors



Infrared Detector Types

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- Cooled detectors
 - Photovoltaic or photoconducting mechanisms; HgCdTe and GaAs quantum well devices
 - High framing rates and low noise
- Uncooled detectors
 - Rely on a thermal response: bolometric, pyroelectric, and thermionic devices
 - Have slow framing rates and are relatively insensitive
 - Lighter and smaller system packaging possible



Uncooled Framing Camera/Pushbroom Pathfinding Thermal Instruments

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- Infrared Spectral Imaging Radiometer (ISIR)
 - Flown on space shuttle mission STS-85 in August 1997 as part of a cloud science experiment
 - Uncooled Lockheed Martin microbolometer array (327x240 pixels)
 - Framing imager with filter wheel
 - 3 narrow bands at 8.55, 10.8, and 11.8 μm
 - 1 broad band at 7-13 μm
 - 250 m GSD
 - 85 km swath from shuttle altitude
 - f-number 0.73, lens diameter 50 mm
 - NEDT 0.01-0.06 $^{\circ}\text{K}$ at all wavelengths with TDIx40 for a 300 $^{\circ}\text{K}$ scene temperature
 - Ambient and cold inflight calibration capability



Uncooled Framing Camera/Pushbroom Pathfinding Thermal Instruments (cont.)

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- Extremely good quality imagery was obtained for each band
- Accuracy goal was met to within a factor of 2 or 3
 - Pre-production prototype detector used

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Potential New Configurations

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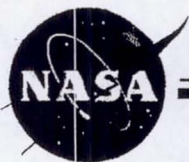
- Cooled detectors
 - Advanced Land Imager (ALI) pushbroom architecture (common telescope for all bands)
 - Multispectral Thermal Imager (MTI)
- Uncooled detectors
 - ALI pushbroom architecture (common telescope for all bands)
 - Custom TIR system separate from ALI
 - Framing camera and filter wheel (ISIR)
 - Pushbroom multispectral system (THEMIS)



Uncooled Thermal Detector Characteristics

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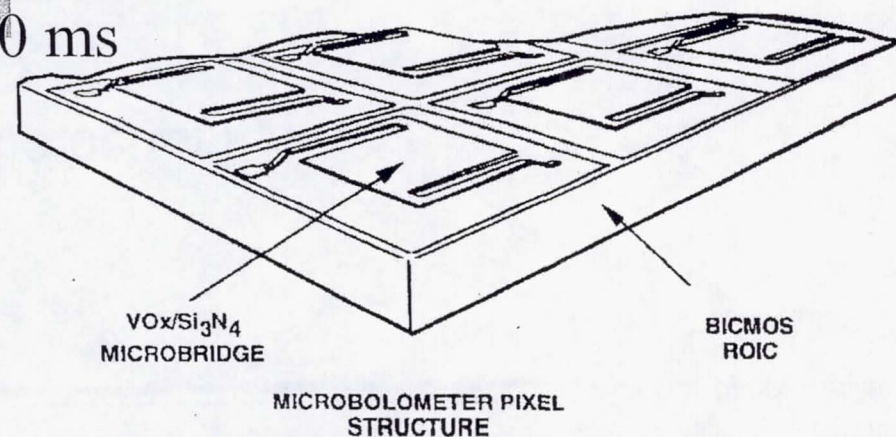
- Primarily developed for military systems:
f/1 optics, 30-60 Hz framing rates
- Microbolometers
 - Silicon micro-machined devices provide excellent thermal isolation from substrate
 - Highest sensitivity demonstrated: f/1 optics, 30-60 Hz framing rate, NEDT ~ 20 mK
- Pyroelectric
 - Older technology: f/1 optics, 30 Hz framing rate, NEDT ~ 100 mK
 - Requires a chopper
- Emerging technologies: Thermionic, SAW Oscillators, Capacitive



Silicon Microbolometers

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- Black absorber with a broadband response (atmospheric window 8-14 microns) thermistor
- Typical FPA
 - 320x240 pixels with 50 μm pitch
 - 640x480 pixels with 25 μm pitch
 - Very high fill factor >90%
 - Thermal time constant ~ 10 ms



W. Radford, D. Murphy, M. Ray, S. Propst, A. Kennedy, J. Kojiro, J. Woolaway, and K. Soch, YEAR. "320 x 240 silicon microbolometer uncooled IRFPAs with on-chip offset correction." SPIE Vol. 2746, pp. 82-92.



Spatial Resolution Drivers

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- Fundamental
 - Orbit altitude (the lower the better)
 - Telescope diameter
 - NEDT
- Engineering
 - Detector size
 - Focal lengths
 - Framing rates
 - Data rates and storage



Thermal GSD Trades

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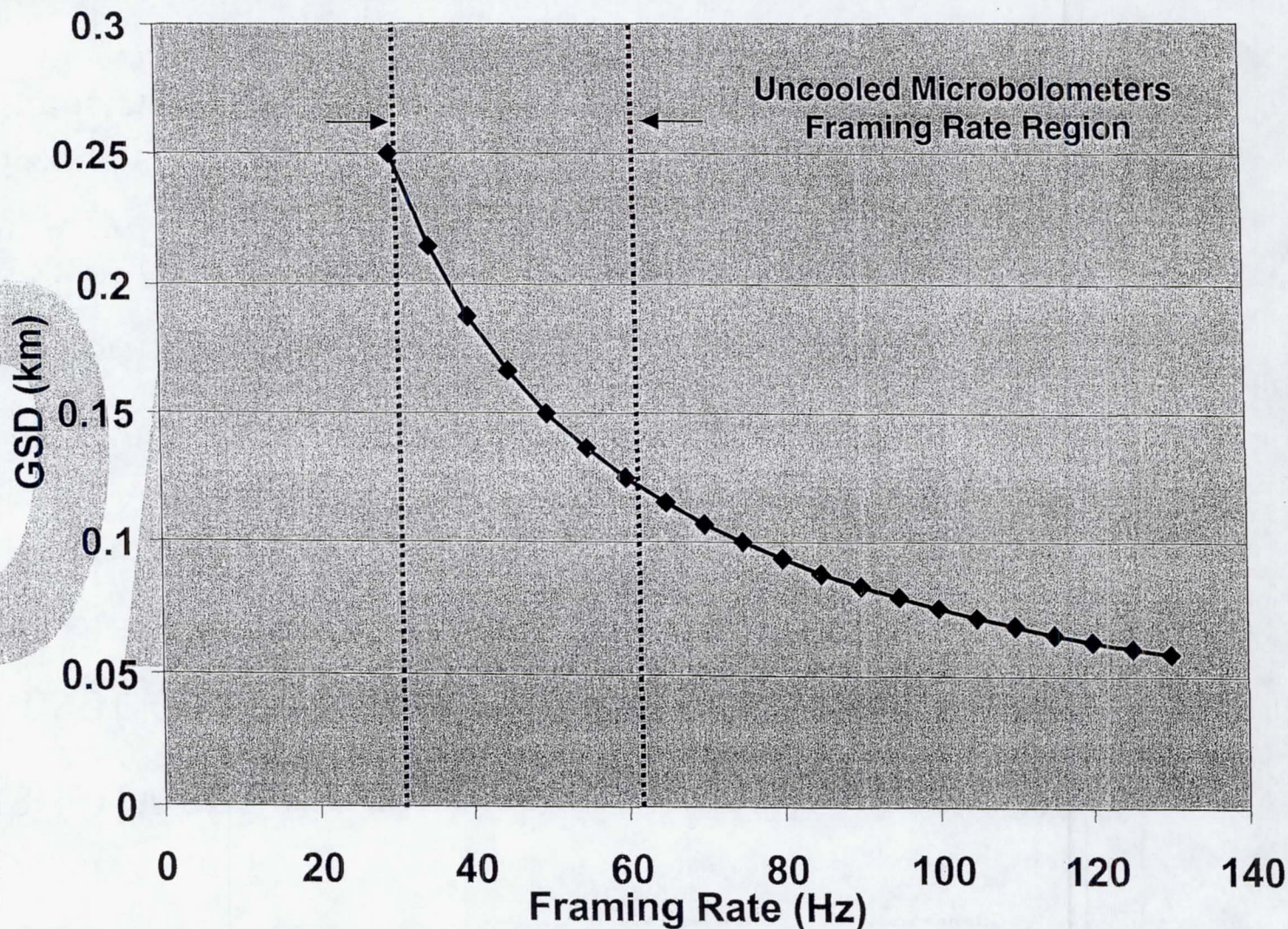
- GSD usually set approximately to FWHM of PSF
- Two sensor characteristics drive PSF:
 - Framing rate
 - Smear is defined as how far a pixel moves in an integration time
 - Typically $GSD \sim \text{Orbital Velocity} / \text{Frame Rate}$
 - Telescope diameter
 - Ground spot size for a diffraction limited system is controlled by Airy Diffraction Pattern
 - Ground spot size $\sim 2.44 \text{ Wavelength} * \text{Range} / \text{Telescope Diameter}$



Frame Rate Trades for Landsat Orbit

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Framing Rate (Hz)	Smear (m)
30.00	250.1
35.00	214.3
40.00	187.5
45.00	166.7
50.00	150.0
55.00	136.4
60.00	125.0
65.00	115.4
70.00	107.2
75.00	100.0
80.00	93.8
85.00	88.3
90.00	83.4
95.00	79.0
100.00	75.0
105.00	71.4
110.00	68.2
115.00	65.2
120.00	62.5
125.00	60.0
130.00	57.7



60 Hz framing rate limits GSD to 125 meters

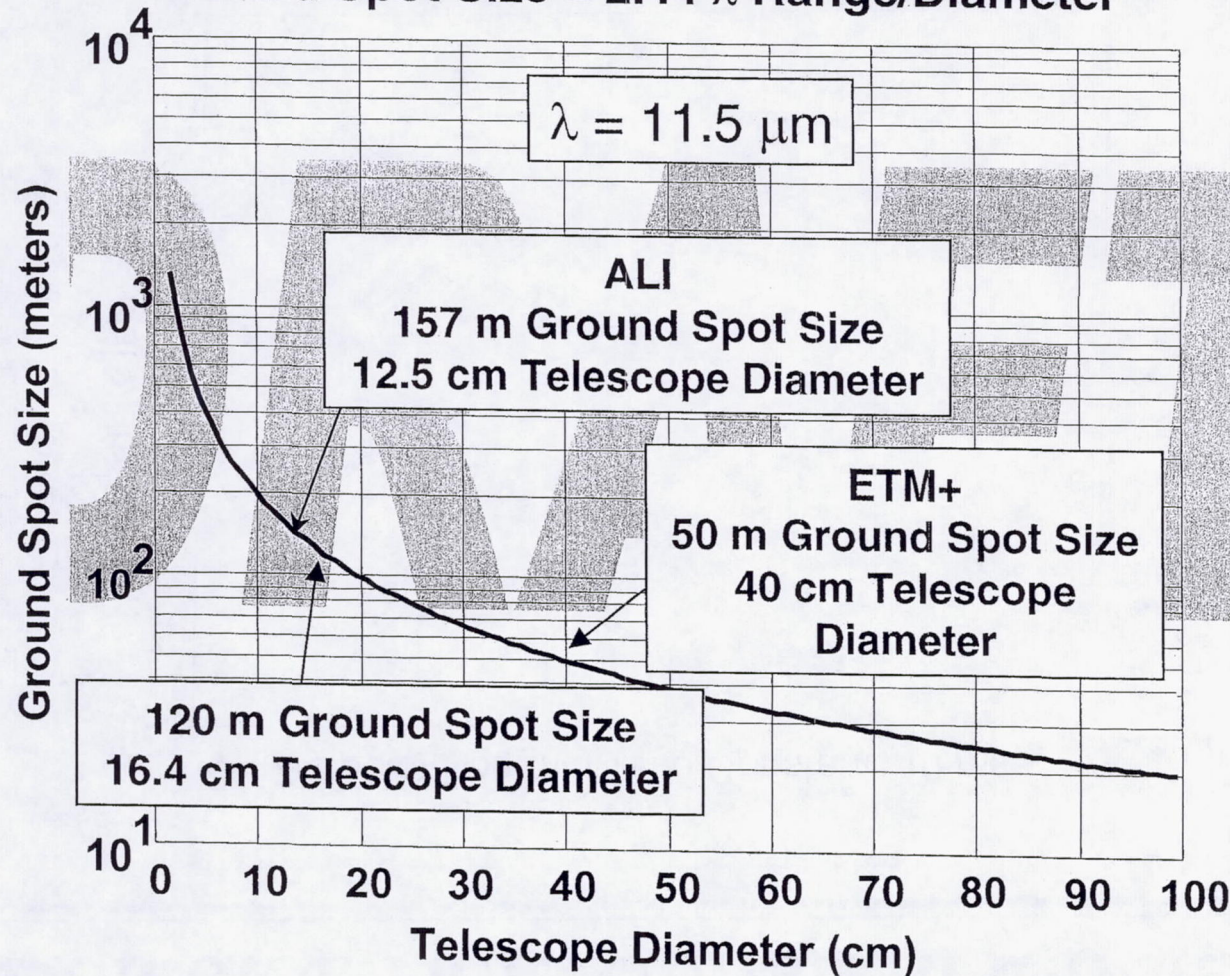


Telescope Diameter Trades

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Diffraction limited resolution (*Rayleigh criteria*)

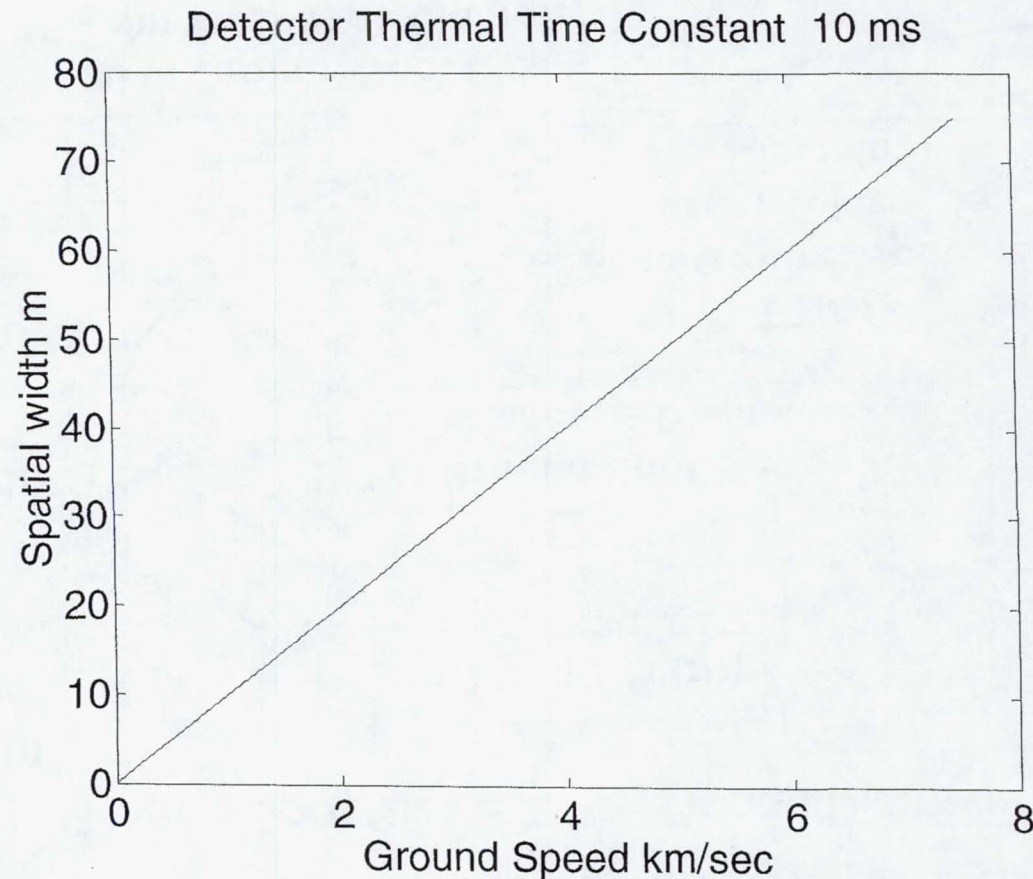
Ground Spot Size = 2.44λ Range/Diameter



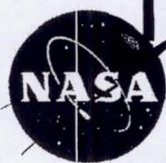


Detector Time Constant Limits

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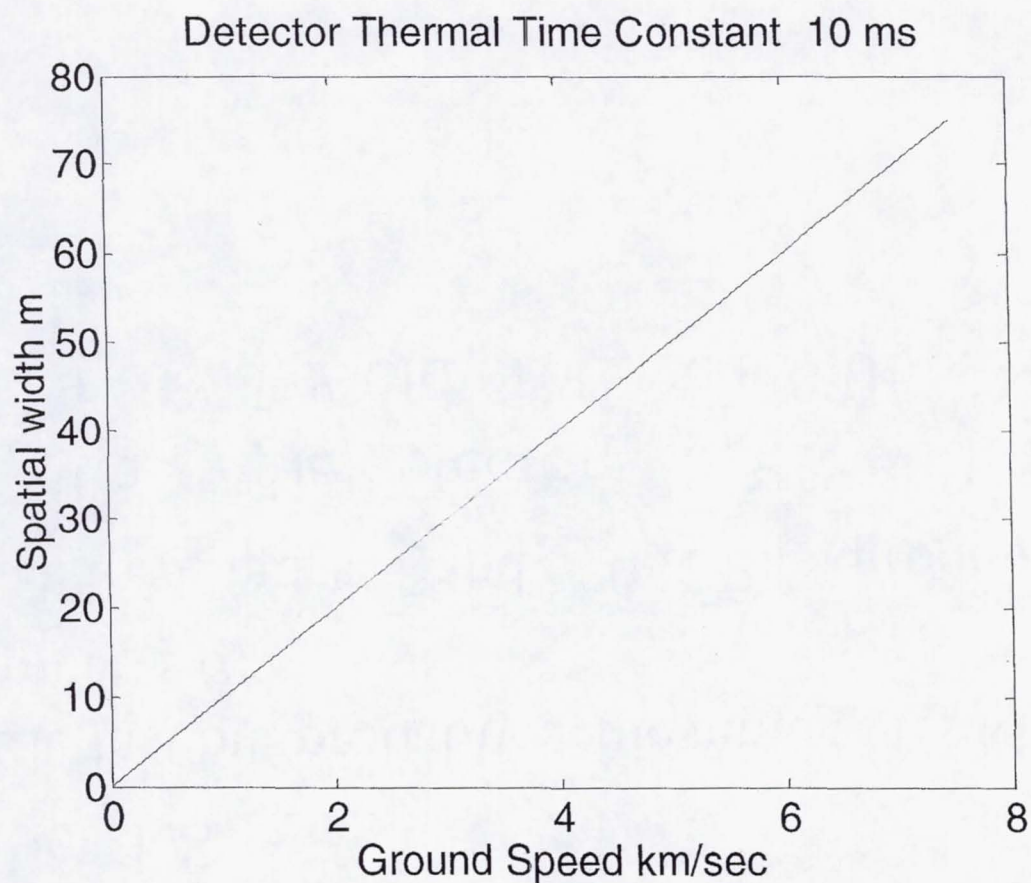


7.5 km/sec produces a 75 m 1/e exponential response



Thermal Time Constant & Spatial Resolution

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Custom Thermal Infrared Sensor (Point Design)

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- F-number 1.8
 - A little fast but probably achievable (THEMIS f-number 1.6)
 - NEDT for Landsat Band 300 K Background
- NEDT of a single detector ~ 0.3 K with a TDI of 1; TDI of 25 could produce a 0.06 k NEDT system!!!!



Custom Thermal Infrared Sensor (Point Design) (cont.)

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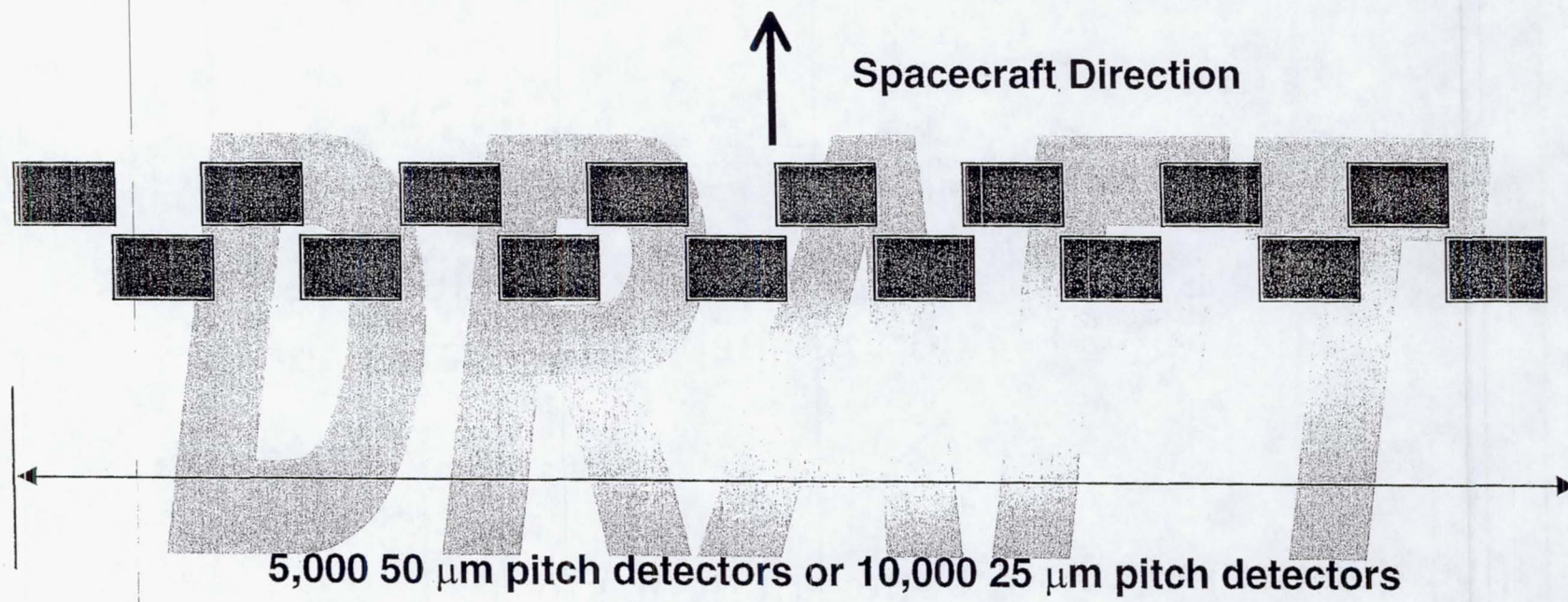
- Detector spot size $50.5 \mu\text{m}$ at $\lambda = 11.5 \mu\text{m}$
 - ~ 1542 detectors at $50 \mu\text{m}$ pitch for 185 km swath with 120 m GSD
 - $\sim$ Five 320×240 FPAs
 - Many detectors available for other bands
 - ASTER-like instrument
 - Oversampling: use $25 \mu\text{m}$ pitch detectors and oversample PSF



Possible FPA Configuration

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Double bank of uncooled microbolometer detector arrays



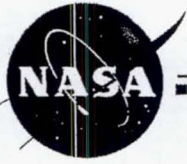
320x240 or 640x480 detector FPA



Backscanning System

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- Trade off coverage for GSD by backscanning the system
 - Used by 1 m class GSD commercial systems such as IKONOS and QuickBird
 - Cutting the ground speed down by a factor of 2 or more produces dramatic results
 - Improved NEDT for a fixed GSD
- 30 m GSD or better system possible with backscanning, creating a high system with MTI-like performance without cooled detectors



Simulation

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- Landsat 7 thermal imagery was synthesized from ATLAS 10-meter imagery
- Microbolometer-simulated imagery was produced with several parameters varied

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LDCM Thermal Trade Simulation

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Microbolometer 30 m detector; 60 m PSF; 125 m smear; 0.3 K NEDT;
decimation (3x3). Final product: 30 m GSD.

L7 Baseline



Microbolometer





LDCM Thermal Trade Simulation

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Landsat 7 60 m detector; 60 m PSF; NEDT 0.22 K; decimation (6x6).

Final product: 60 m GSD.



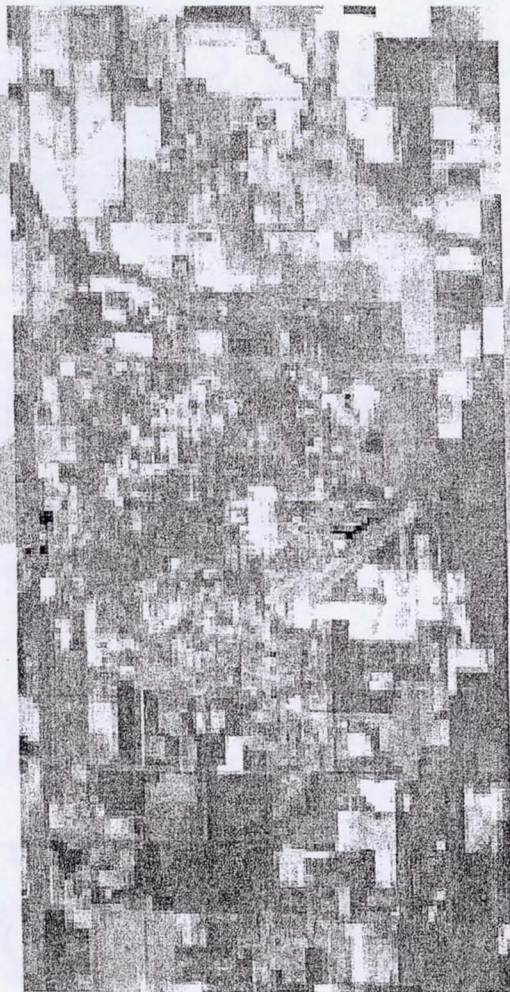


LDCM Thermal Trade Simulation

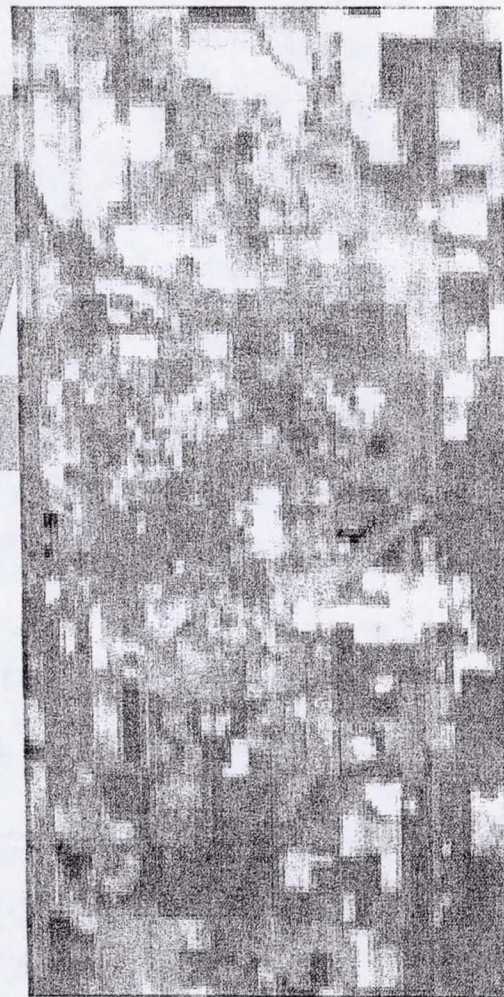
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Microbolometer 60 m detector; 120 m PSF; 125 m smear; NEDT 0.30 K;
LPF (2x2); decimation (6x6). Final product: 60 m GSD.

L7 Baseline



Microbolometer





LDCM Thermal Trade Simulation

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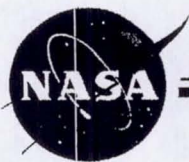
Microbolometer 60 m detector; 120 m PSF; 125 m smear; no noise; LPF (2x2); decimation (6x6). Final product: 60 m GSD.

L7 Baseline



Microbolometer





LDCM Thermal Trade Simulation

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Microbolometer 30 m detector; 60 m PSF; 125 m smear; no noise; decimation (3x3). Final product: 30 m GSD.

L7 Baseline



Microbolometer





Summary

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- Conventional microbolometer pushbroom mode offers potential for low-cost LDCM thermal or ASTER capability with at least a 60-120 meter GSD
- Backscanning could produce MTI-like performance without cooled detectors
- Cooled detectors could produce hyperspectral thermal class system or extremely high spatial resolution class instrument

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